

System size dependence of particle production in EPOS and some remarks about low energies

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in collaboration with

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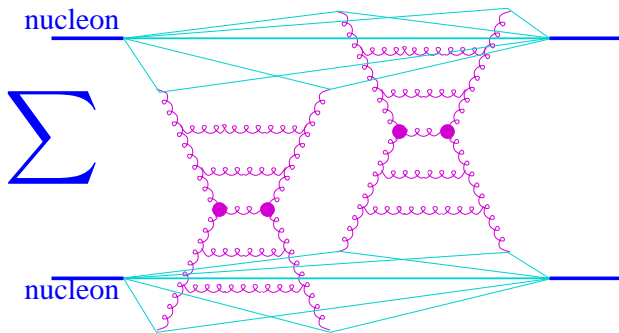
EPOS = Gribov-Regge approach

S-Matrix based on Pomerons

Pomerons :
Parton ladders (initial and final state radiation, DGLAP)

Cutting rules to get inelastic cross sections.

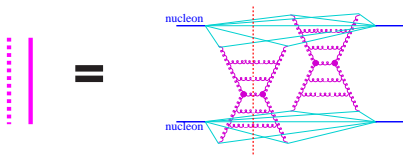
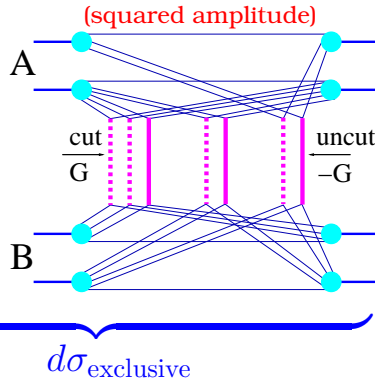
Same principle for pp, pA, AA



Explicite formulas for cross sections

(even partial cross sections)

$$\sigma^{\text{tot}} = \sum_{\text{cut } P} \int \sum_{\text{uncut } P} \int$$



=> kinky strings

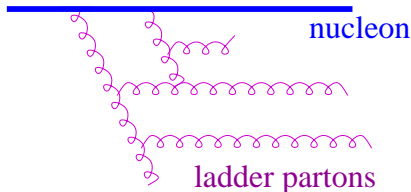


Non-linear effects

Computing the expressions G for single Pomeron:
A cutoff Q_0 is needed (for the DGLAP integrals).

**Taking Q_0 constant leads to a power law increase
of cross sections vs energy (=> wrong)**

because non-linear effects like
gluon fusion are not taken into
account



Solution: Instead of a constant Q_0 , use a dynamical **saturation scale for each Pomeron:**

$$Q_s = Q_s(N_{\text{IP}}, s_{\text{IP}})$$

with

- N_{IP} = **number of Pomerons connected to a given Pomeron (whose probability distr. depends on Q_s)**
- s_{IP} = **energy of considered Pomeron**

We get $Q_s(N_{\text{IP}}, s_{\text{IP}})$ from fitting

- the energy dependence of elementary quantities ($\sigma_{\text{tot}}, \sigma_{\text{el}}, \sigma_{\text{SD}}, dn^{\text{ch}}/d\eta(0)$) for pp
- the multiplicity dependence of dn^{π}/dp_t at large p_t for pp at 7 TeV

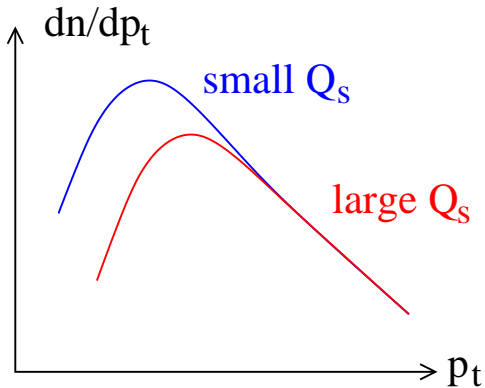
We find

$$Q_s \propto \sqrt{N_{\text{IP}}} \times (s_{\text{IP}})^{0.30}$$

CGC for AA:

$$Q_s \propto N_{\text{part}} \times (1/x)^{0.30}$$

Parton distributions



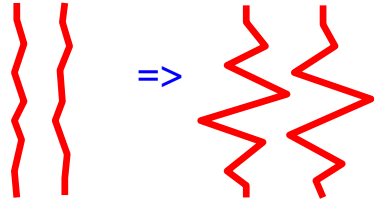
Increasing multiplicity

=> increasing N_{Pom}

=> Increasing Q_s

=> harder Pomerons

=> harder strings



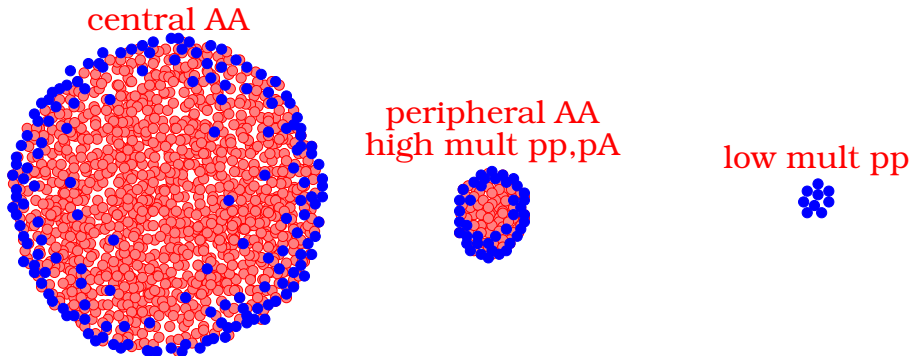
=> more high p_t particles

=> Strong increase of $\langle p_t \rangle$ with multiplicity

and gives a strong nonlinear increase of D or J/Psi multiplicity vs charged multiplicity in pp and pPb ...

Core-corona picture in EPOS

Gribov-Regge approach => (Many) kinky strings
=> core/corona separation (based on string segments)



core => hydro => flow + statistical decay
corona => string decay

EPOS status and perspective:

Status 2015: Two parallel developments

EPOS LHC:

Gribov Regge approach, parameterized flow as in EPOS1.99, tuned to LHC data (2012), **very much used (and tested) by LHC pp groups, UE, forward physics etc, and used for air shower simulations**

EPOS 3.0xx:

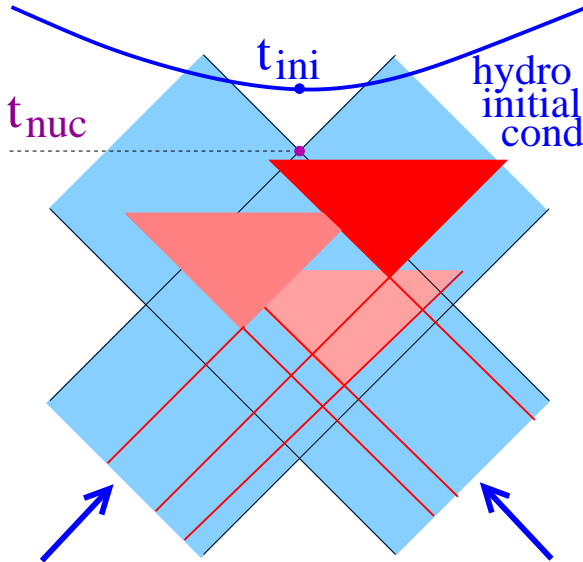
Gribov Regge approach, viscous hydro, parton saturation, **mainly used for HI and collectivity in pp**

2015/2016/2017: “Fusion”, to accommodate basic pp and HI features, public version;

Currently: EPOS3.2xx (beta version)

What about EPOS at low energies?

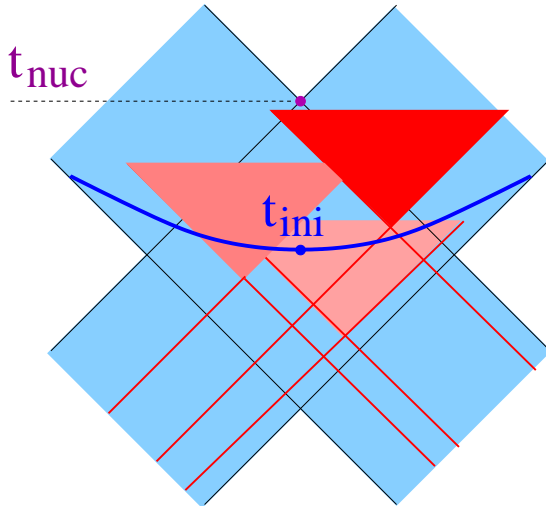
Space time picture of a HI collision



longitudinal dimension of each nucleus: $2R/\gamma$

$t_{\text{nuc}} = R/\gamma v$
big at low E

**too little core,
too little flow for
 $t_{\text{ini}} > t_{\text{nuc}}$**



Solution:

take

$$t_{ini} < t_{nuc}$$

Back to LHC:

Testing EPOS 3.210

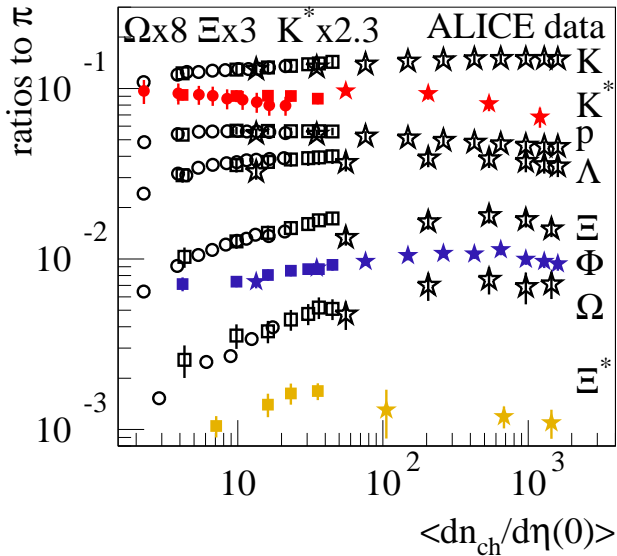
To get a global overview:

□ **Chemistry:** Particle ratios vs $\left\langle \frac{dn_{\text{ch}}}{d\eta}(0) \right\rangle$
for pp, pPb, PbPb

□ **Flow:** Average transverse momenta vs $\left\langle \frac{dn_{\text{ch}}}{d\eta}(0) \right\rangle$
for pp, pPb, PbPb

$\left\langle \frac{dn_{\text{ch}}}{d\eta}(0) \right\rangle$ for multiplicity classes defined via forw multiplicities

Particle ratios to pions vs $\left\langle \frac{dn_{ch}}{d\eta}(0) \right\rangle$



circles = pp (7TeV)

squares = pPb (5TeV)

stars = PbPb (2.76TeV)

ALICE data references (collected by A. G. Knospe)

$\langle dN_{ch}/d\eta \rangle$ in Pb+Pb: Phys. Rev. Lett. 106 032301 (2011)

π^+ , K^+ , p^+ in Pb+Pb: Phys. Rev. C 88 044910 (2013)

Lambda in Pb+Pb: Phys. Rev. Lett. 111 222301 (2013)

Ξ and Omega in p+Pb: Phys. Lett. B 758 389-401 (2016)

π^+ , K^+ , p^+ , Λ in p+Pb: Phys. Lett. B 728 25-38 (2014)

$\langle dN_{ch}/d\eta \rangle$ in p+Pb: Eur. Phys. J. C 76 245 (2016)

Ξ and Omega in p+Pb: Phys. Lett. B 758 389-401 (2016)

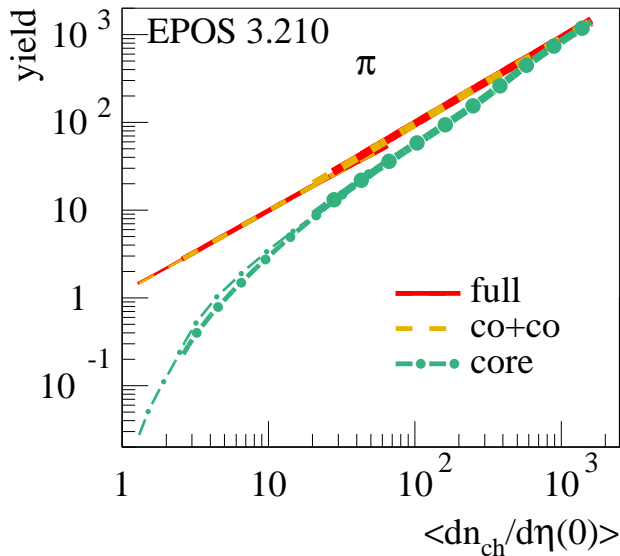
$\langle dN_{ch}/d\eta \rangle$ p+p 7 TeV: Eur. Phys. J. C 68 345-354 (2010)

π^+ , K^+ , p^+ in p+p 7 TeV: Eur. Phys. J. C 75 226 (2015)

Ξ and Omega in p+p 7 TeV: Phys. Lett. B 712 309 (2012)

and pp data points from Rafael Derradi de Souza, SQM2016

Pion yields: core / corona contribution



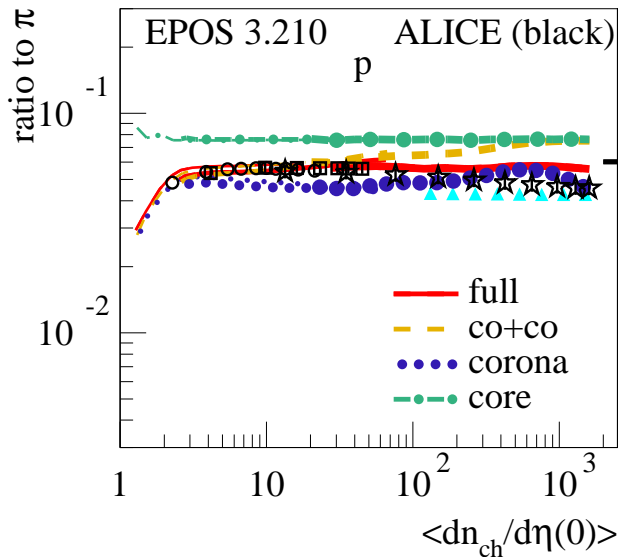
thin lines
= pp (7TeV)

intermediate lines
= pPb (5TeV)

thick lines
= PbPb (2.76TeV)

full = with hadronic
cascade (UrQMD)

Proton to pion ratio



core hadronization:

$T = 164 \text{ MeV}, \mu_B = 0$

statistical model fit

(horizontal black line)

A. Andronic et al.,

arXiv:1611.01347

$T = 156.5 \text{ MeV}, \mu_B = 0.7 \text{ MeV}$

thin lines = pp (7TeV)

intermediate lines = pPb (5TeV)

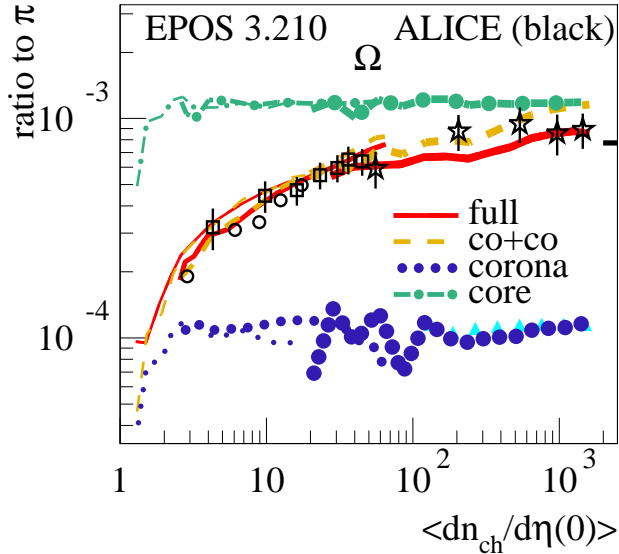
thick lines = PbPb (2.76TeV)

circles = pp (7TeV)

squares = pPb (5TeV)

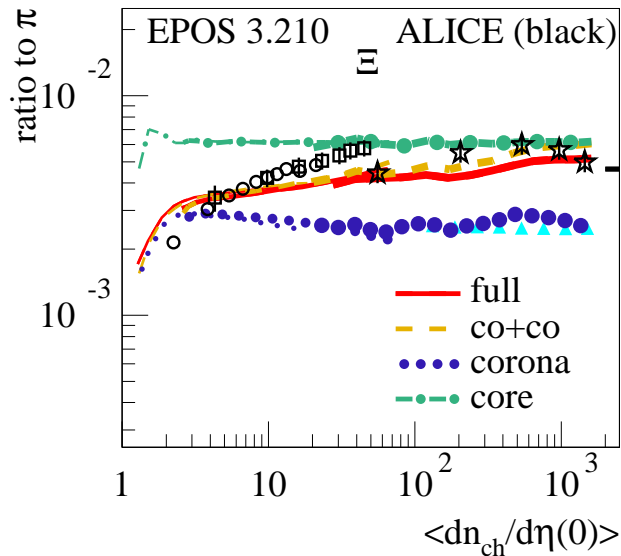
stars = PbPb (2.76TeV)

Omega to pion ratio

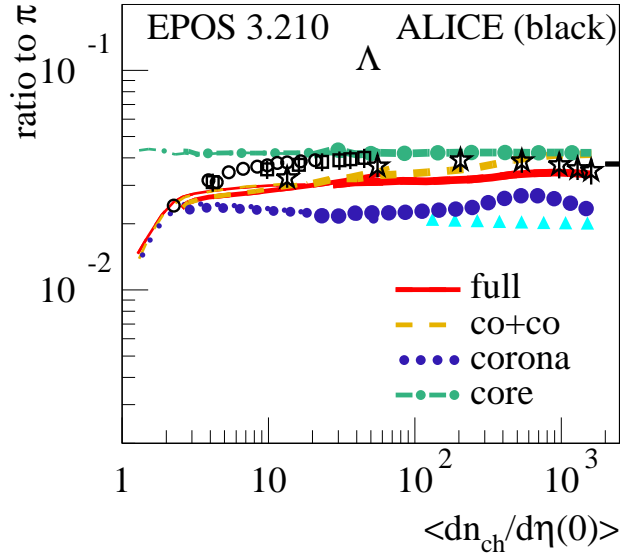


thin lines = pp (7TeV)
 intermediate lines = pPb (5TeV)
 thick lines = PbPb (2.76TeV)
 circles = pp (7TeV)
 squares = pPb (5TeV)
 stars = PbPb (2.76TeV)

Xi to pion ratio

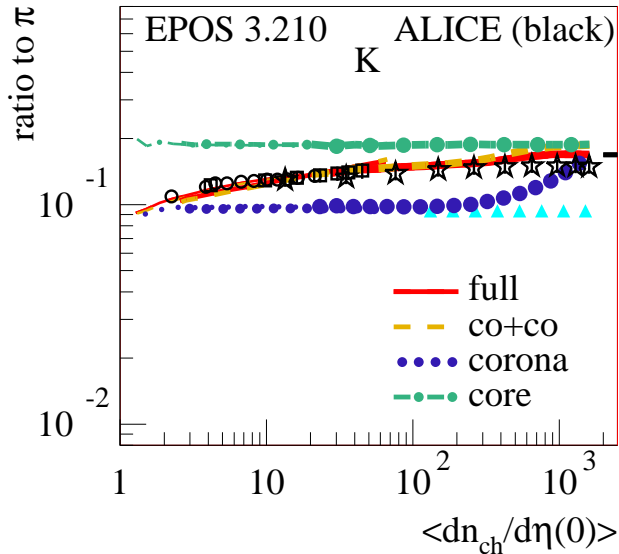


Lambda to pion ratio



thin lines = pp (7TeV)
 intermediate lines = pPb (5TeV)
 thick lines = PbPb (2.76TeV)
 circles = pp (7TeV)
 squares = pPb (5TeV)
 stars = PbPb (2.76TeV)

Kaon to pion ratio



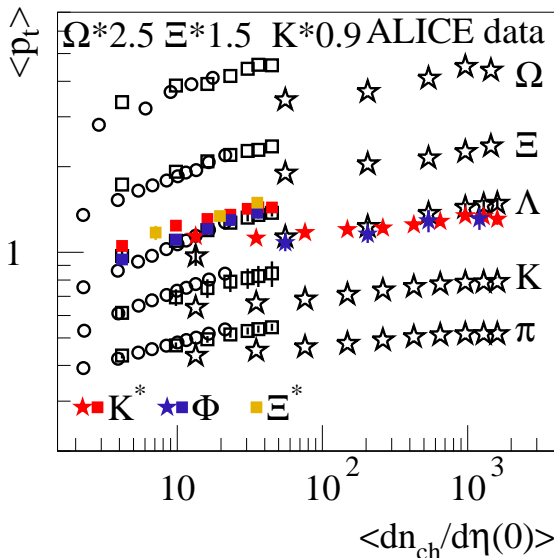
Ratios h/π for $h = p, K, \Lambda, \Xi, \Omega$ vs $\left\langle \frac{dn}{d\eta}(0) \right\rangle$:

Core and corona contributions separately roughly constant

Difference (core - corona) increasing for $p \rightarrow K, \Lambda \rightarrow \Xi \rightarrow \Omega$

=> increasing slope

Mean p_t vs $\left\langle \frac{dn_{ch}}{d\eta}(0) \right\rangle$

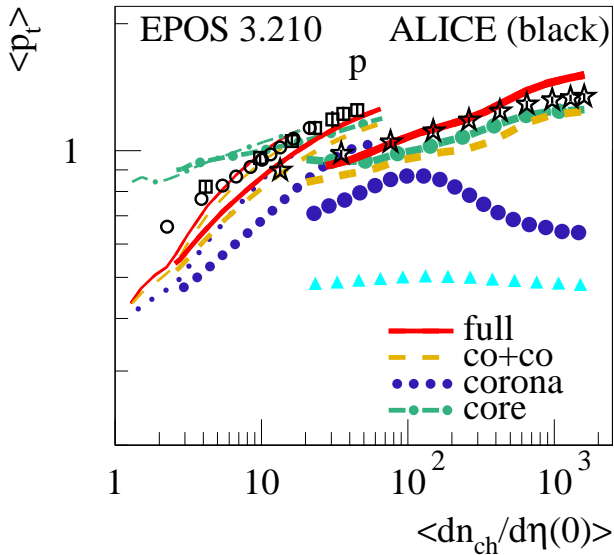


circles = pp (7TeV)

squares = pPb (5TeV)

stars = PbPb (2.76TeV)

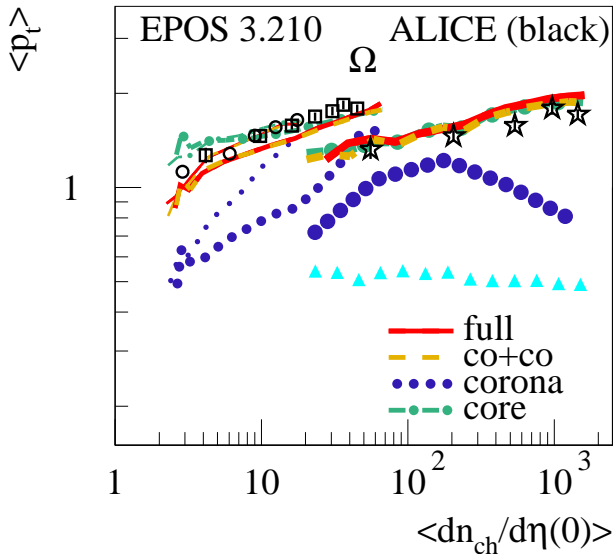
Average p_t of protons



thin lines = pp (7TeV)
 intermediate lines = pPb (5TeV)
 thick lines = PbPb (2.76TeVV)

circles = pp (7TeV)
 squares = pPb (5TeV)
 stars = PbPb (2.76TeV)

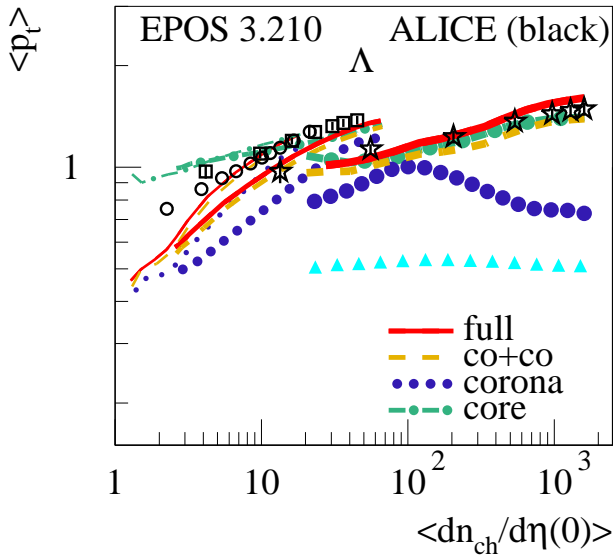
Average p_t of Omegas



thin lines = pp (7TeV)
 intermediate lines = pPb (5TeV)
 thick lines = PbPb (2.76TeVV)

circles = pp (7TeV)
 squares = pPb (5TeV)
 stars = PbPb (2.76TeV)

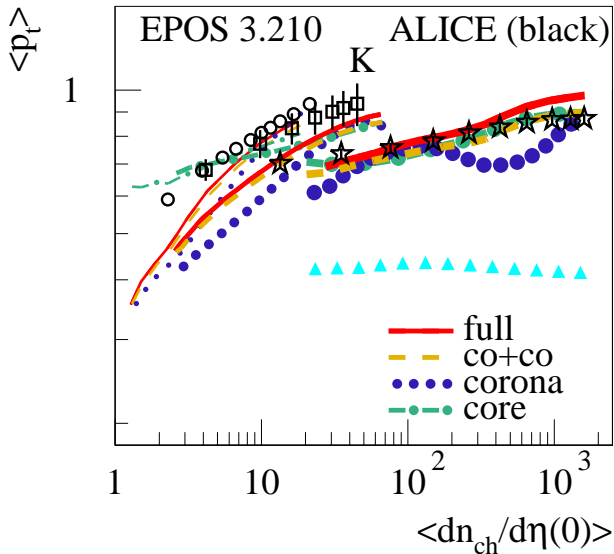
Average p_t of lambdas



thin lines = pp (7TeV)
 intermediate lines = pPb (5TeV)
 thick lines = PbPb (2.76TeVV)

circles = pp (7TeV)
 squares = pPb (5TeV)
 stars = PbPb (2.76TeV)

Average p_t of kaons



thin lines = pp (7TeV)
 intermediate lines = pPb (5TeV)
 thick lines = PbPb (2.76TeV)
 circles = pp (7TeV)
 squares = pPb (5TeV)
 stars = PbPb (2.76TeV)

Average p_t of $K, p, \Lambda, \Xi, \Omega$ vs $\left\langle \frac{dn}{d\eta} \right\rangle_{(0)}$:

Moderate increase of core contribution
(same for pp and pPb, similar to PbPb)

Strong increase of corona contribution
(stronger for pp than for pPb, much stronger than for PbPb)

Slope(pp) > slope(pPb) >> slope(PbPb)

K, π : pp-pPb splitting

The multiplicity dependence of the corona contribution is crucial

Summary

- **To understand multiplicity dependence of particle production we have to understand the corona contribution (= non-flow).**
- **The latter one dominates low multiplicity pp, but its relative weight decreases continuously with multiplicity (but is never zero)**
- **Investigating the multiplicity dependence of particle ratios and mean pt in pp, pA, AA: EPOS's core-corona picture describes the trend**
- **Strong increase of corona pt due to the $N_{P_{om}}$ dependence of the saturation scale ...**

Core => Hydro evolution (Yuri Karpenko)

Israel-Stewart formulation, $\eta - \tau$ coordinates, $\eta/S = 0.08$, $\zeta/S = 0$

$$\partial_{;\nu} T^{\mu\nu} = \partial_\nu T^{\mu\nu} + \Gamma_{\nu\lambda}^\mu T^{\nu\lambda} + \Gamma_{\nu\lambda}^\nu T^{\mu\lambda} = 0$$

$$\gamma (\partial_t + v_i \partial_i) \pi^{\mu\nu} = -\frac{\pi^{\mu\nu} - \pi_{\text{NS}}^{\mu\nu}}{\tau_\pi} + I_\pi^{\mu\nu} \quad \gamma (\partial_t + v_i \partial_i) \Pi = -\frac{\Pi - \Pi_{\text{NS}}}{\tau_\Pi} + I_\Pi$$

$T^{\mu\nu} = \epsilon u^\mu u^\nu - (p + \Pi) \Delta^{\mu\nu} + \pi^{\mu\nu}$,

$\pi_{\text{NS}}^{\mu\nu} = \eta (\Delta^{\mu\lambda} \partial_{;\lambda} u^\nu + \Delta^{\nu\lambda} \partial_{;\lambda} u^\mu) - \frac{2}{3} \eta \Delta^{\mu\nu} \partial_{;\lambda} u^\lambda$

$\partial_{;\nu}$ denotes a covariant derivative,

$\Pi_{\text{NS}} = -\zeta \partial_{;\lambda} u^\lambda$

$\Delta^{\mu\nu} = g^{\mu\nu} - u^\mu u^\nu$ is the projector orthogonal to u^μ ,

$I_\pi^{\mu\nu} = -\frac{4}{3} \pi^{\mu\nu} \partial_{;\gamma} u^\gamma - [u^\nu \pi^{\mu\beta} + u^\mu \pi^{\nu\beta}] u^\lambda \partial_{;\lambda} u_\beta$

$\pi^{\mu\nu}$, Π shear stress tensor, bulk pressure

$I_\Pi = -\frac{4}{3} \Pi \partial_{;\gamma} u^\gamma$

Freeze out: at 164 MeV, Cooper-Frye $E \frac{dn}{d^3p} = \int d\Sigma_\mu p^\mu f(up)$, equilibrium distr

Hadronic afterburner: UrQMD

Marcus Bleicher, Jan Steinheimer